

This question paper contains 4+2 printed pages]

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S. No. of Question Paper : 873

Unique Paper Code : 222602 G

Name of the Paper : Statistical Physics [PHHT-620]

Name of the Course : B.Sc. (Hons.) Physics

Semester : VI

Duration : 3 Hours

Maximum Marks : 75

*(Write your Roll No. on the top immediately on receipt of this question paper.)*

Attempt five questions in all.

Question No. 1 is compulsory.

All questions carry equal marks.

Symbols have their usual meanings.

I. Answer any five of the following :

5×3=15

(a) Explain the significance of partition function in statistical thermodynamics.

P.T.O.

- (b) Calculate the wavelength corresponding to maximum emission from the Sun's surface at a temperature of 6000 K. ( $b = 2898 \mu\text{m K}$ )
- (c) Discuss the limitations of law of equipartition of energy.
- (d) Under what conditions do the Bose-Einstein and Fermi-Dirac distribution approach the Maxwell-Boltzmann distribution.
- (e) Give the equilibrium temperatures of two systems at temperature +300 K and -300 K in thermal contact.
- (f) Write two properties of photons which make them different from other bosons.
- (g) The Fermi energy for metal A is 3.15 eV. Find its value for metal B given that the free electron density in metal B is nine times that in metal A.
- (h) What are the basic assumptions of Planck's theory of Black body radiation ?

2. (a) Explain the meaning of thermodynamic probability. How does thermodynamic probability differ from conventional definition of probability ? 3
- (b) Establish the relation between entropy and thermodynamic probability and show that the constant occurring in the relation is the Boltzmann's constant. 7.5
3. (a) Prove that the single particle partition function for an ideal monoatomic gas enclosed in a cube (volume V, maintained at temperature T) is given by :

$$Z_s = \left( \frac{2\pi mk_B T}{h^2} \right)^{3/2} V$$

where,  $m$  is the mass of each particle of gas and  $h$  is the Planck's constant. 4

- (b) Assuming that this gas consists of  $N$  identical and indistinguishable particles, prove that the expression for the entropy ( $S$ ) of this gas is given by :

$$S = Nk_B \left( \frac{5}{2} + \ln \frac{Z_s}{N} \right)$$

and hence show that this expression resolves Gibbs' paradox. 7.4

4. (a) State and prove Kirchoff's law for Black body radiation.

Discuss one of its applications. 13,1

- (b) Derive the expression for pressure exerted by diffuse radiation. 5

- (c) Show thermodynamically that when the radiations enclosed in an enclosure are adiabatically expanded,  $TV^{1/3}$  is a constant. 5

5. (a) Show that the number of modes of vibrations per unit volume of an enclosure in the frequency range  $\nu$  and  $\nu + d\nu$  is given by :

$$N_\nu d\nu = \frac{8\pi\nu^2}{c^3} d\nu.$$

Using this expression, deduce the Rayleigh-Jean's law of energy distribution. 8,2

- (b) Prove that the energy emitted per unit area per second from a black body is proportional to the fourth power of its absolute temperature. 5

6. (a) Show that number of microstates associated with a given macrostate for B-E statistics is given by :

$$W = \prod_i \frac{(n_i + g_i - 1)!}{n_i! (g_i - 1)!}$$

where,  $n_i$  represents the number of particles in the energy level  $\epsilon_i$  having degeneracy  $g_i$ . 4

- (b) Using Stirling's approximation and maximizing  $S = k_B \ln(W)$ , subject to the constraints that total energy  $E$  and total number of particles  $N$  are constant, show that B-E distribution function is given by :

$$n_i = \frac{g_i}{e^{\alpha + \beta\epsilon_i} - 1}$$

where,  $\alpha$  and  $\beta$  are Lagrange multipliers. 8

- (c) Show that chemical potential of a boson gas is negative. 3

7. (a) Obtain the expressions for  $E_{F0}$ , the Fermi energy at  $T = 0$  K and  $P_0$ , the zero-point pressure of electron gas. 5,2

- (b) Prove that the average kinetic energy per particle of Fermi gas at  $T = 0$  K is given by :

$$\langle E \rangle = \frac{3}{5} E_{F_0} \quad 6$$

- (c) Why  $\langle E \rangle$  in part (b) is not zero ? 2
8. (a) Show that the matter in white dwarf stars behave like a strongly degenerate relativistic electron gas. 5
- (b) What is Bose-Einstein condensation ? Derive an expression for the temperature at which this phenomenon occurs. 28